

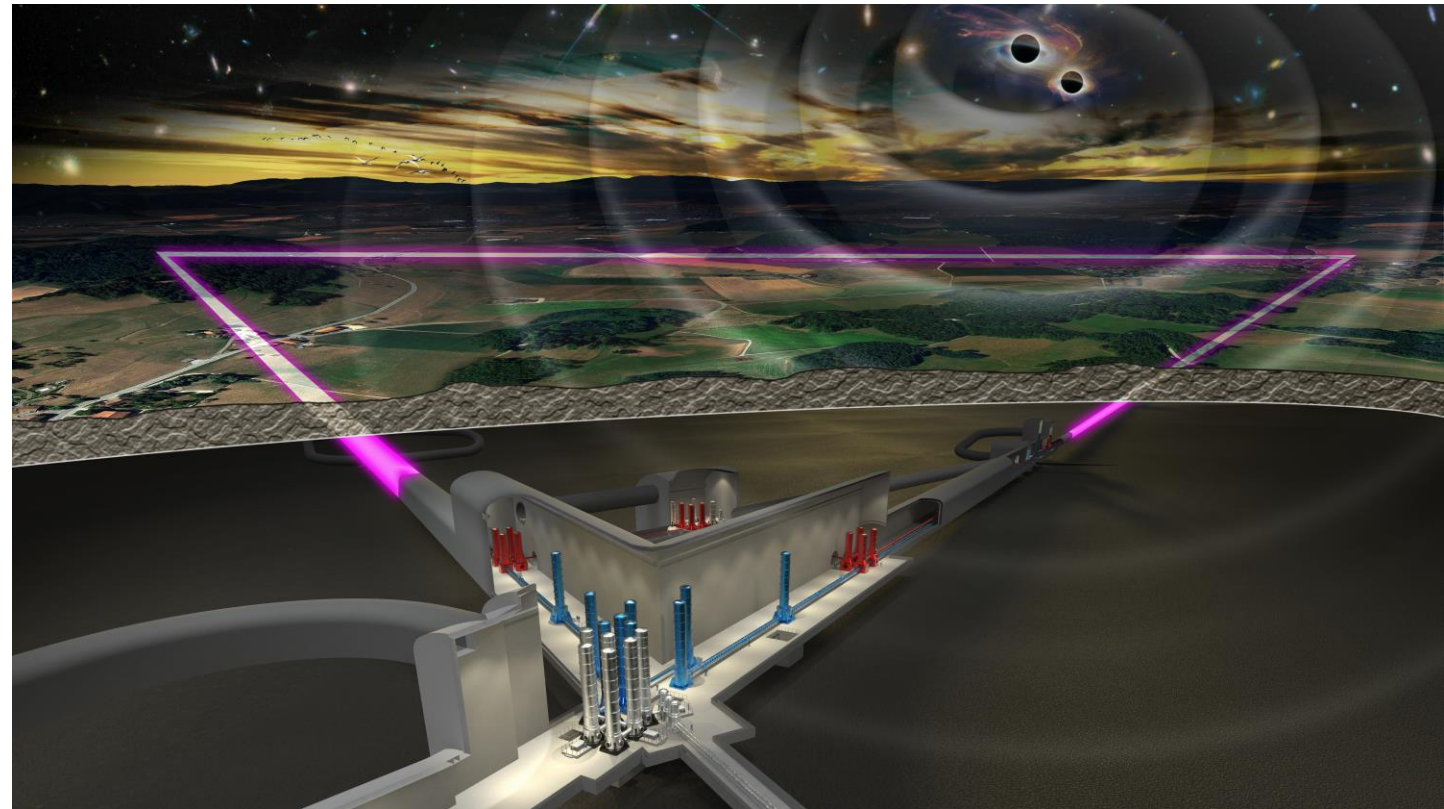
The Science of the Einstein Telescope

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The Einstein Telescope is a future observatory of gravitational waves

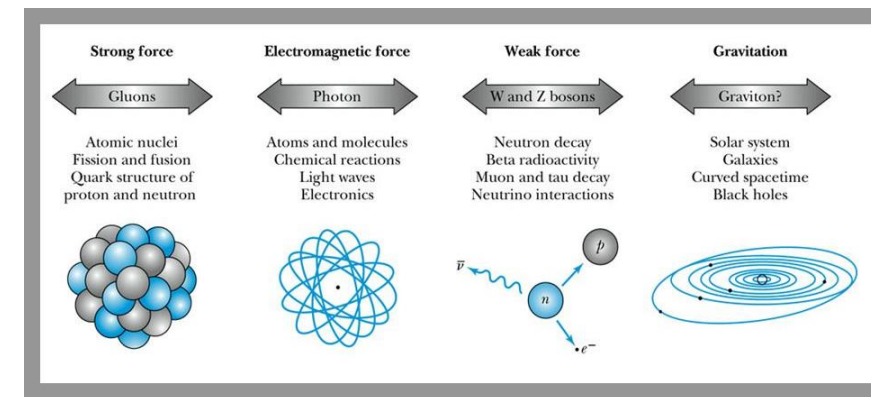
- Gravity
- Universe



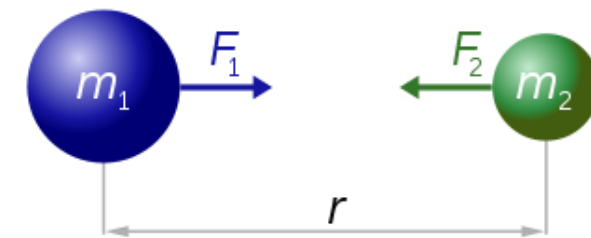
Gravity

- One of the four fundamental interactions

- The only one not yet described by a **quantum** relativistic theory



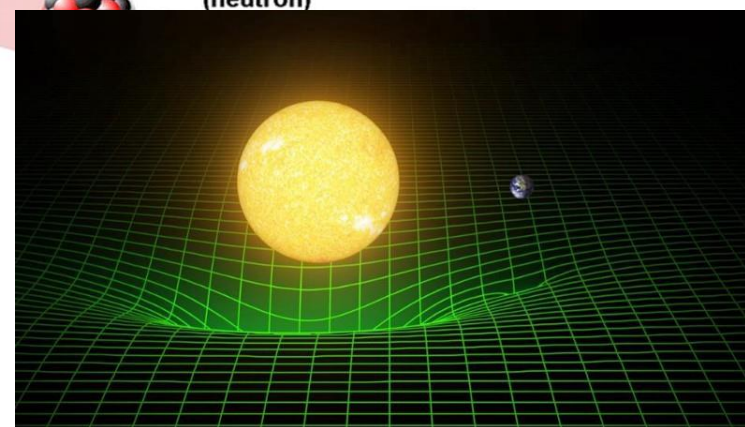
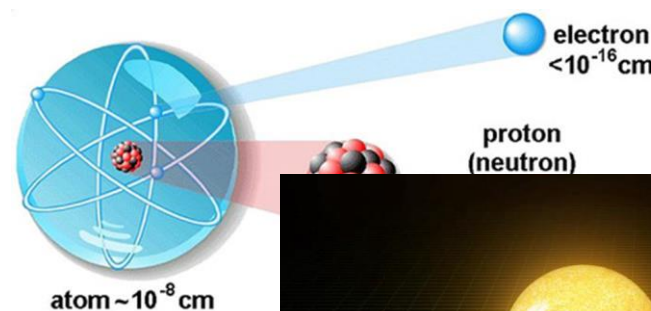
- 1687 Newton's theory



$$F_1 = F_2 = G \frac{m_1 \times m_2}{r^2}$$

- 1915: Einstein theory of General Relativity

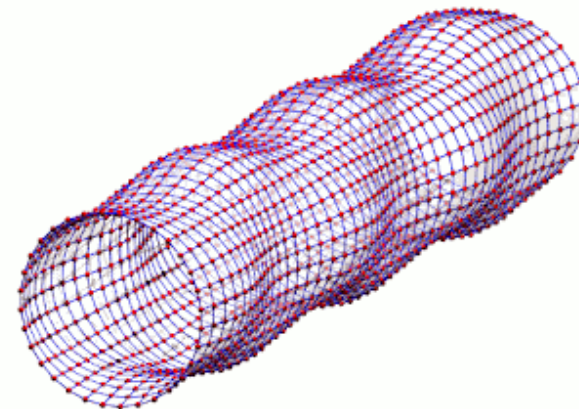
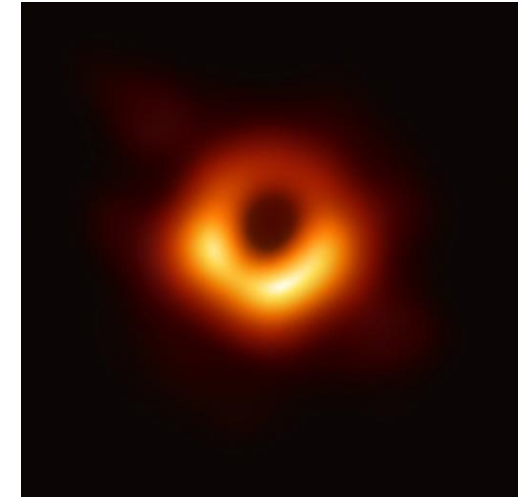
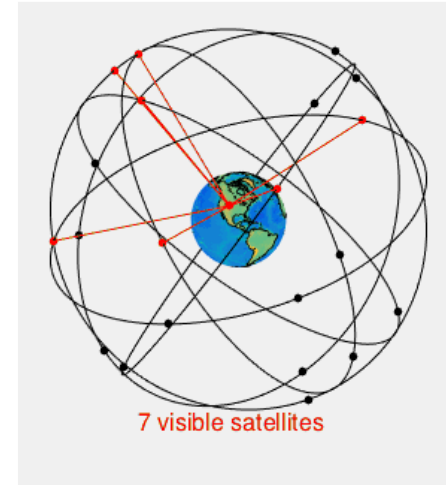
- Gravity is related to space, time and matter



$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

Some predictions of General Relativity

- Space and time are « curved » by matter
 - e.g. atomic clocks on GPS satellites run as fast as $\sim 40 \mu\text{s}$ per day w.r.t. those on earth
- Black holes
- Gravitational waves
 - Space-time deformations travelling at the speed of light
 - Observed for the first time in 2015



The universe

- Universe observations so far via:

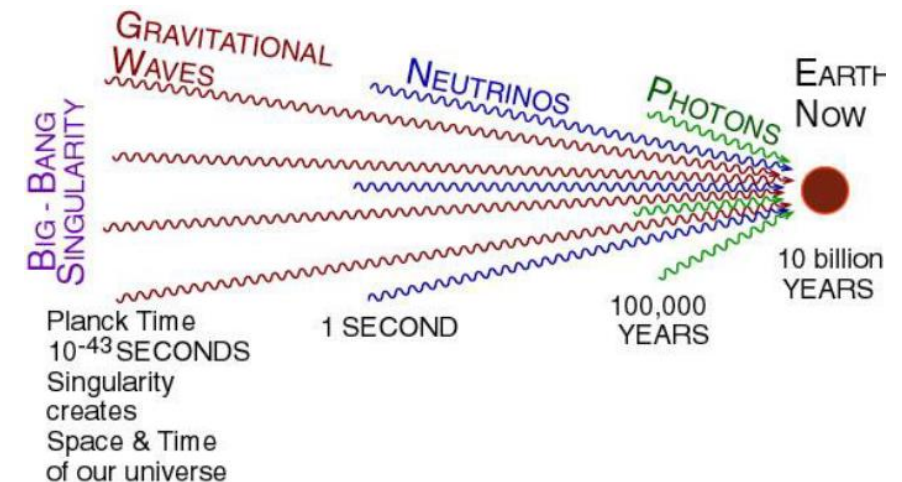
- photons (e.m. waves, x-rays, γ -rays)
- cosmic rays (p, e, nuclei)
- neutrinos

- GW revolution: unperturbed, rich, far-reaching information

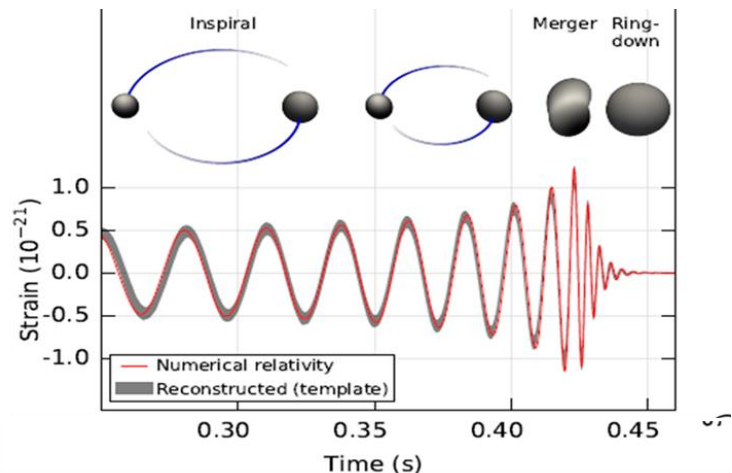
- Astrophysics
- fundamental interactions (gravity, strong interaction)
- Cosmology (dark matter, accelerating universe expansion, universe evolution)



Image credit: NAOJ/ALMA
<http://alma.mtk.nao.ac.jp/>



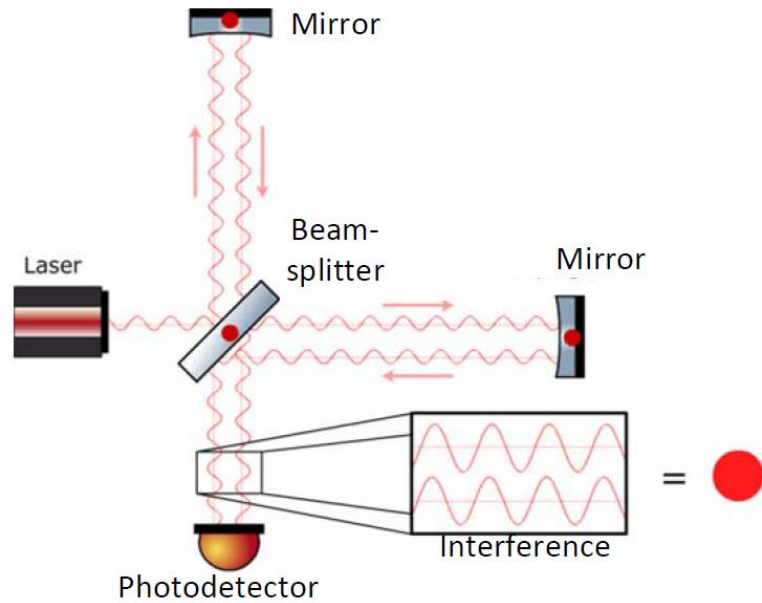
- First direct detection by LIGO/Virgo collaborations in 2015
 - Nobel prize 2017 for R. Weiss, K. Thorne and B. Barish
 - 2016 G. Lemaitre prize from UCLouvain to K. Thorne



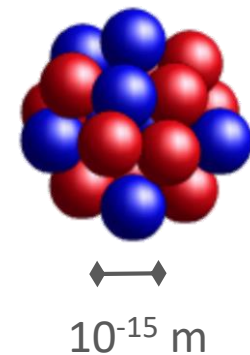
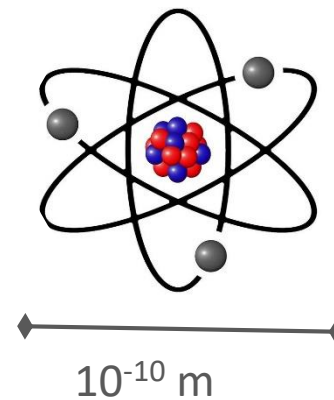
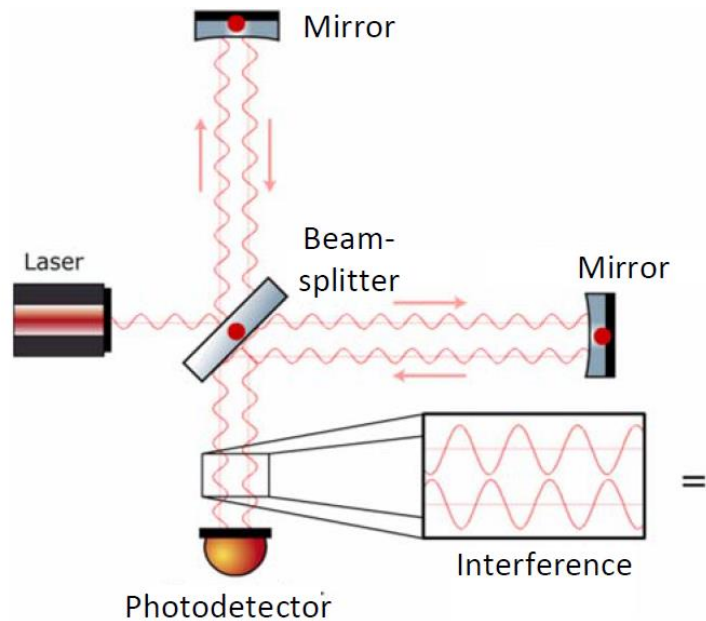
Le Prix Georges Lemaître à Kip Thorne



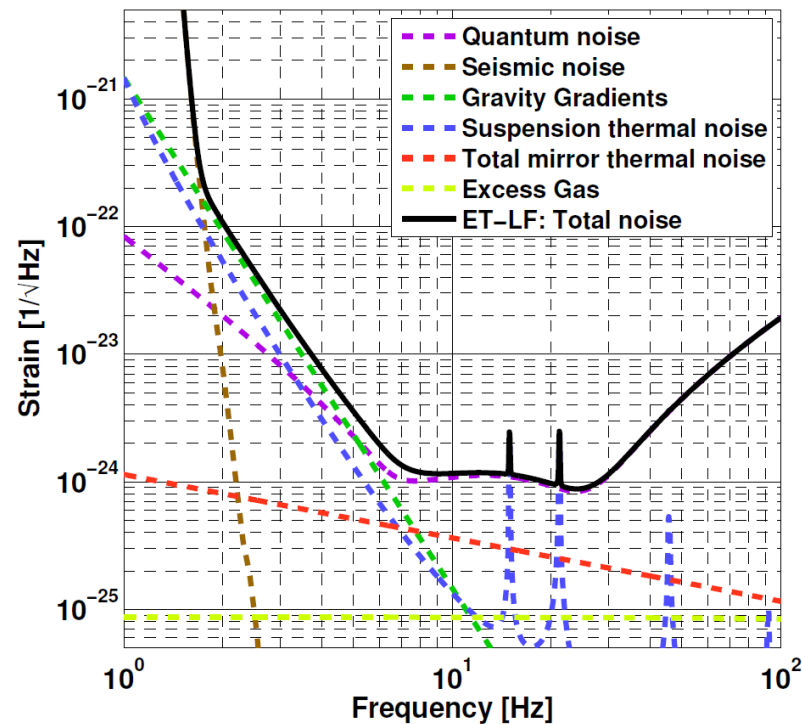
GW detection principle



- Laser interferometers measure changes in arm length difference
- Effect of merger of two BHs (30 solar masses each) at 10^9 light years $\sim 10^{-18}$ m
- Laser wavelength $\sim 10^{-6}$ m



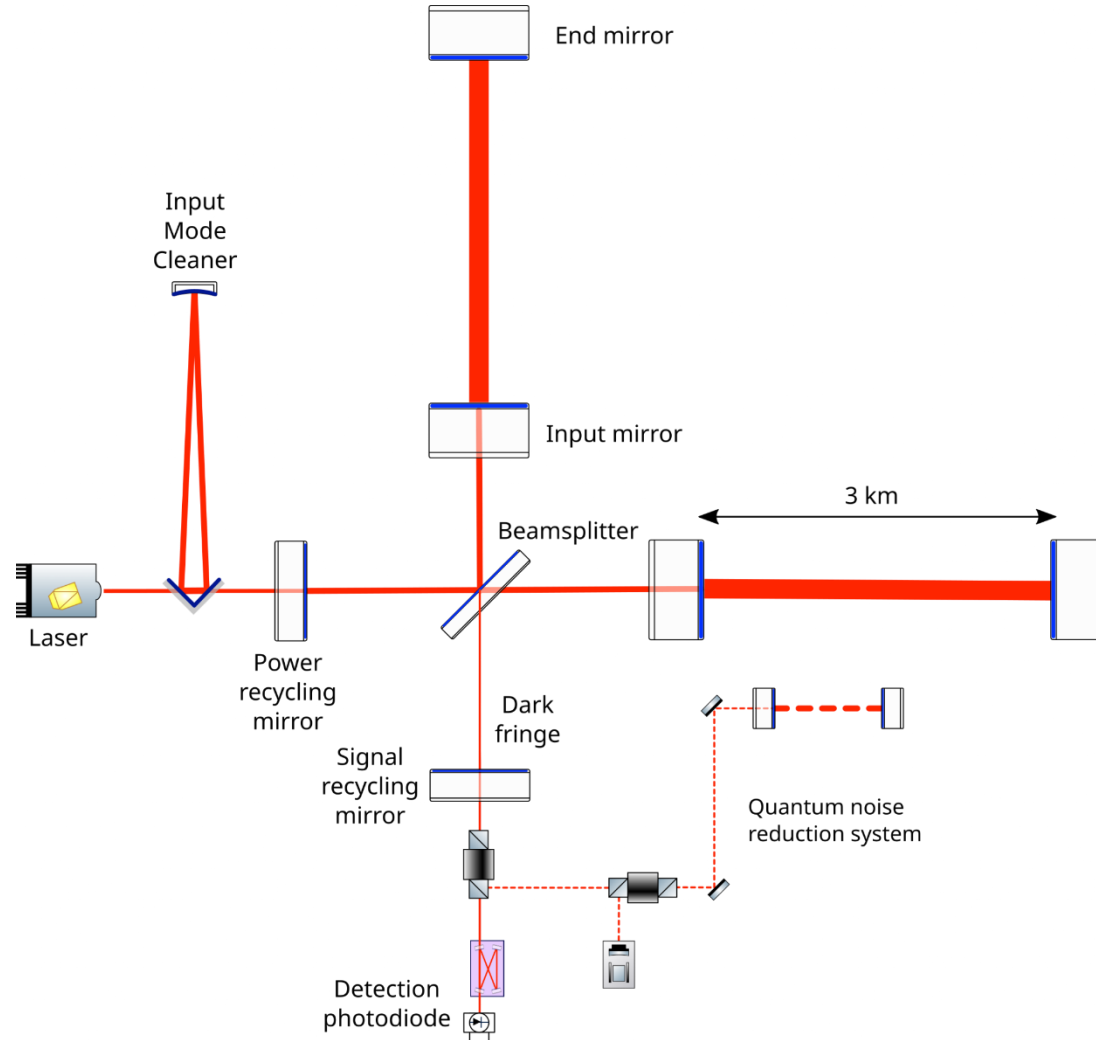
Detector sensitivity



- Seismic noise
- Gravity gradient noise
- Thermal noise
- Quantum noise
- Excess gas

.. and myriad of technical noise sources

A GW observatory in operation (Virgo)



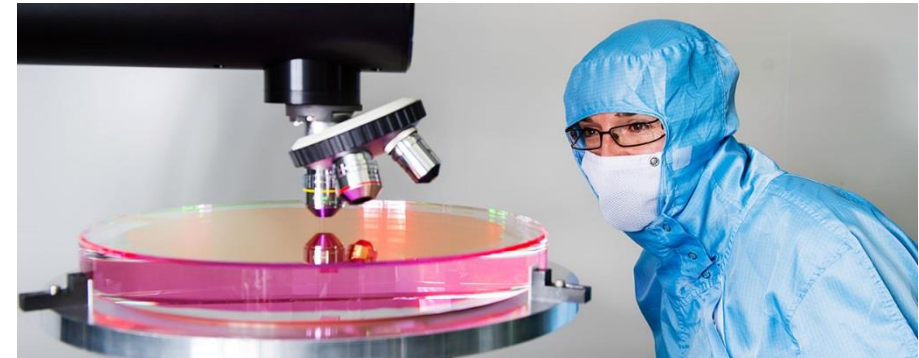
GW Technology

Pushing frontiers in mechatronics, lasers and optics, material sciences, vacuum and cryogeny, controls, ...

Measuring and attenuating vibrations



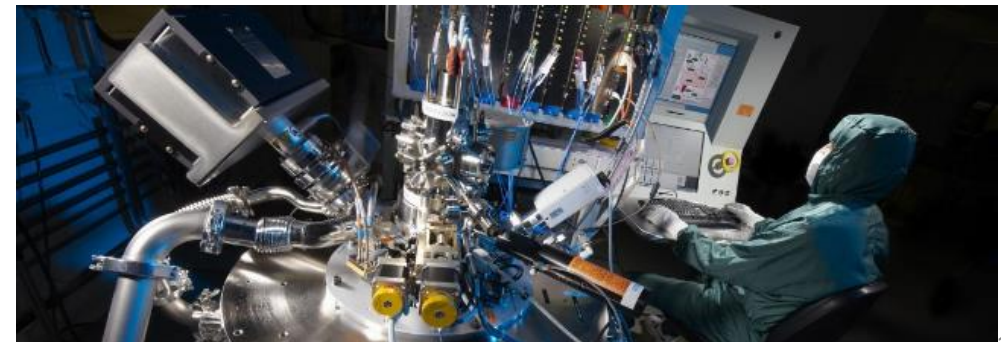
Optics, coatings, special materials, laser technology, semiconductor technology



Vacuum technology



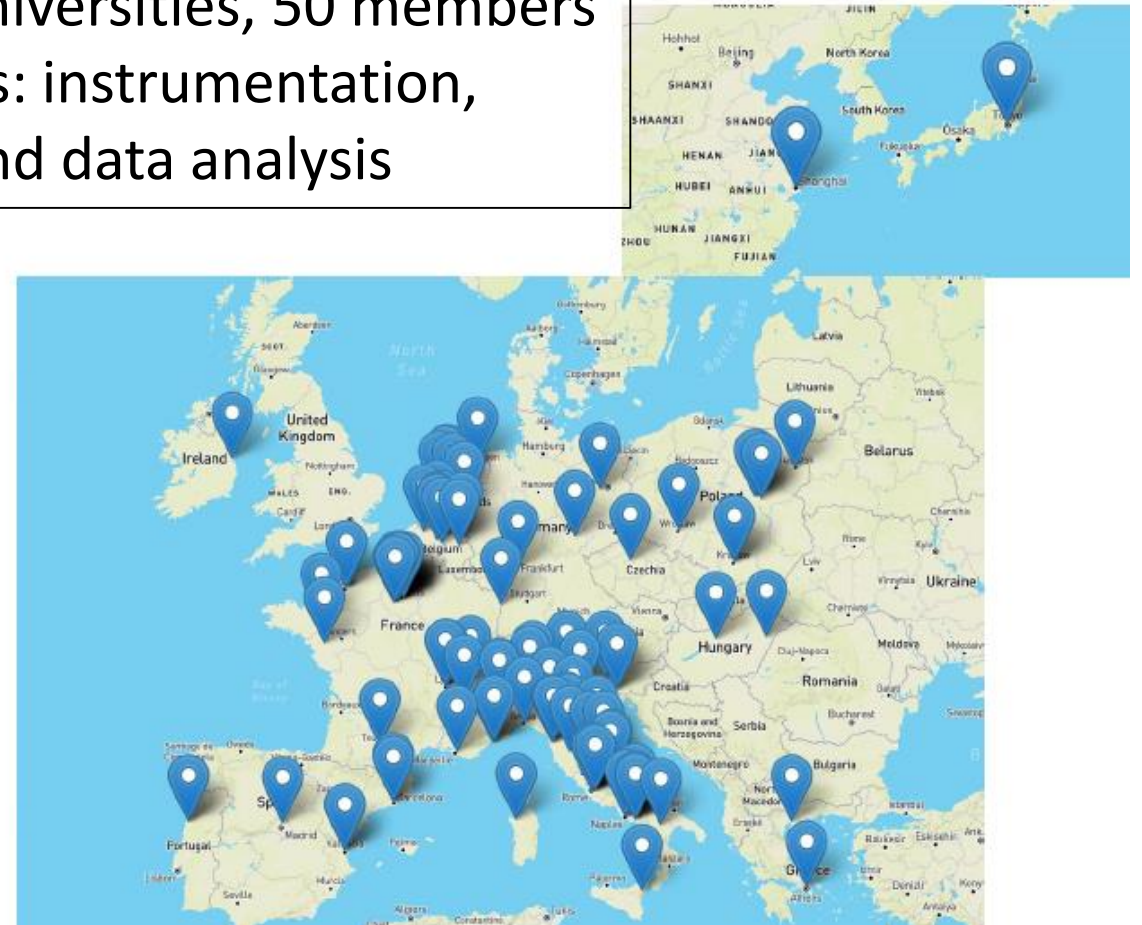
Cryogenic systems



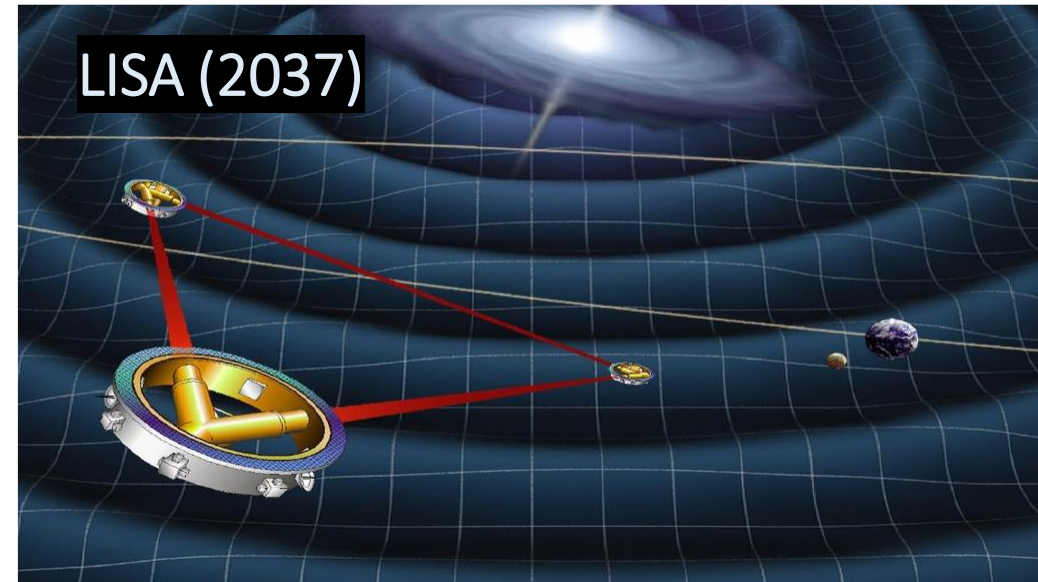
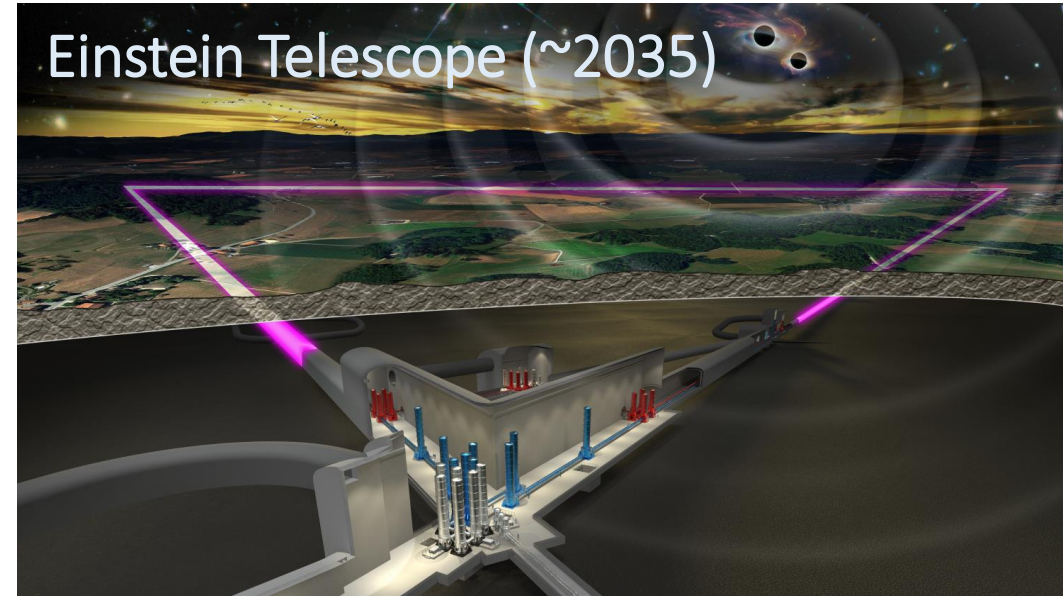
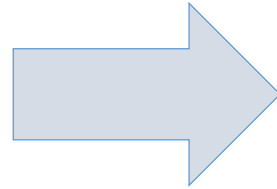
Virgo collaboration

- ~800 members, ~450 authors, 136 institutions from 15 countries
- 36 Groups:
 - 32 full members
 - 4 in the first year
- 9 countries

- Belgium: 7 universities, 50 members
- Contributions: instrumentation, computing and data analysis



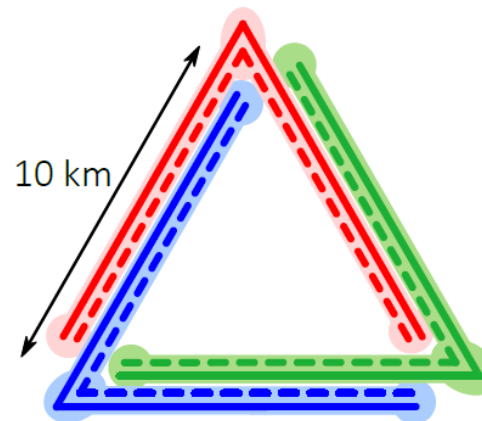
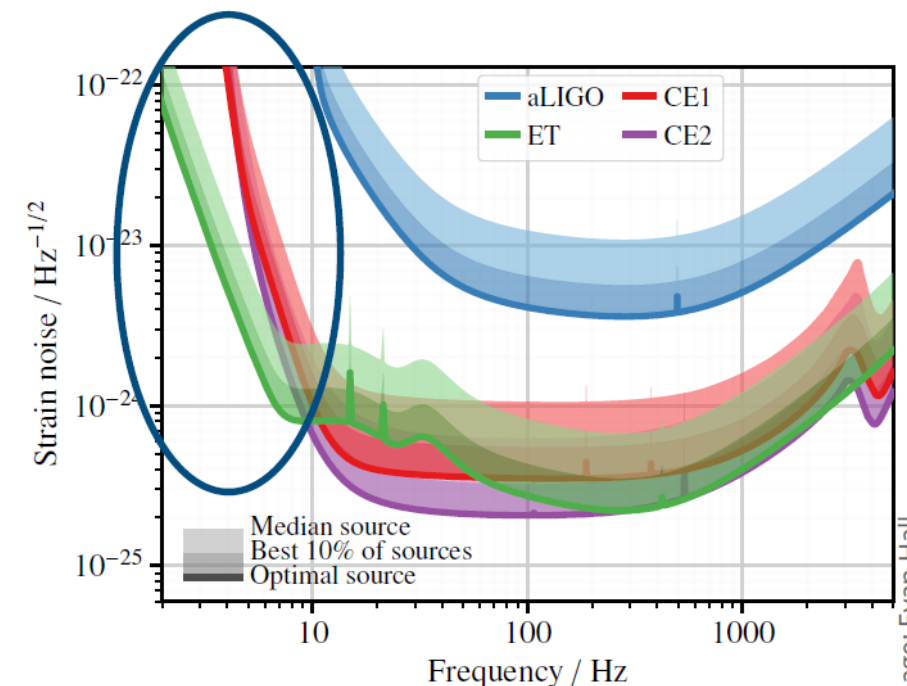
From current detectors to the next generation

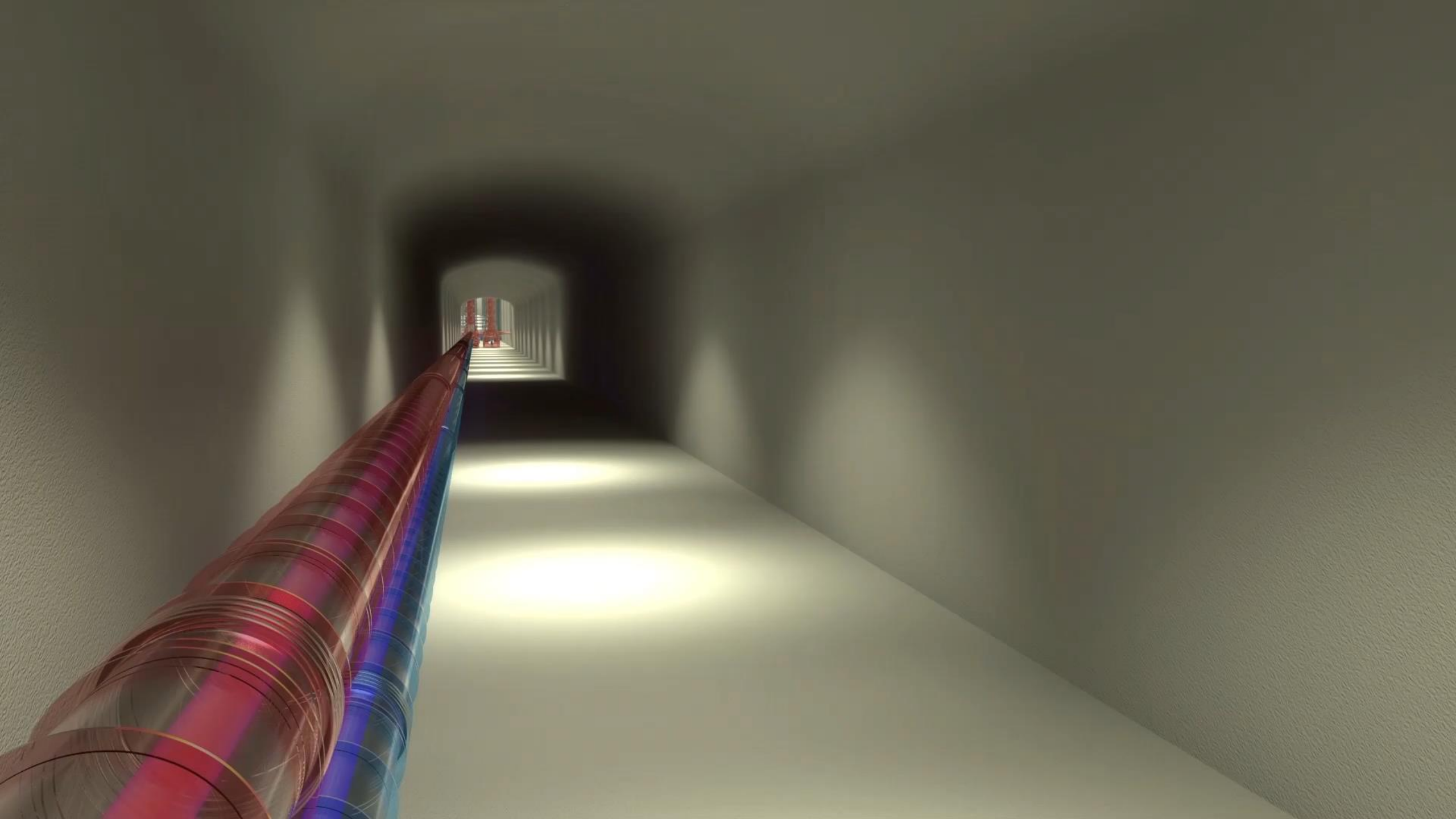


ein Telescope

Einstein Telescope design

- Longer arms (10 km)
 - Effect of GW is equivalent to a relative change in the arm lengths ($\Delta L/L$)
 - All mirror displacement noise sources reduced by making L larger
 - but** this increases laser beam size (need for large mirrors)
- Underground operation (~250 m)
 - Reduction of seismic and gravity gradient noise
 - key for sensitivity at low frequency
- Triangular shape
 - Wave polarizations
 - Null streams
 - Redundancy
 - Sky coverage
 - Single compact infrastructure
- “Xylophone design”





ET site

EMR site

Sos
Ennatos site



• Science

- Seismic, Newtonian, anthropogenic noise

• Cost and feasibility

- Geology, topography, climate, access, services, local regulations

• Socio-political factors

- Availability of funds

Preparing the ground for ET

Research and development facilities for ET technology in the Euregio Meuse-Rhine

ETpathfinder



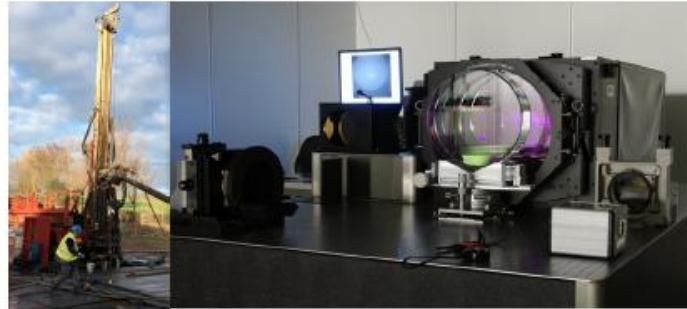
Objective: Development of a model infrastructure for testing new gravitational wave detector technologies and concepts in a complete interferometer in an ET-like environment

Location: UMaastricht-NL

Budget: € 14,8 million

Duration : 2019 – 2022

E-TEST



Objective: Development of ET-technology

- Geological exploration of the EMR and determination of the optimal ET location.
- Development of advanced prototypes for cryogenics, optics and seismic isolation.

Location: CSL ULiège - BE

Budget: € 15,0 million

Duration : 2020 – 2023

ET2SMEs



Objective: Promotion of cooperation between SMEs, large companies and R&D institutions that deal with ET-relevant key technologies in a broad understanding and towards multiple application fields by initiating SME-driven cross-border R&D projects.

Budget: € 2,23 million

Duration : 2021 – 2023

BACK-UP

System	GW amplitude
Tennis ball rotating around a post attached to a 1 m string and at 10 m distance	$h \sim 1. 10^{-54}$
Hulse and Taylor pulsar	$h \sim 1. 10^{-26}$
Io around Jupiter	$h \sim 2. 10^{-25}$
Binary neutron star merger at the Milky Way center	$h \sim 2. 10^{-19}$
BH (30 Msun) binary merger at 10^9 parsecs	$h \sim 1. 10^{-21}$

GW detectors

Living Rev. Relativity, 14, (2011), 5

